

- 1) Conservation laws. For each of the following reactions, (a) establish whether it is allowed or not; (b) if it is not, give the reasons (there may be more than one); (c) establish the interaction that allows it.

$$p + \bar{p} \rightarrow \Lambda^{\circ} + \Lambda^{\circ}$$

$$p \rightarrow n + e^{+}$$

$$\mu^{+} \rightarrow \nu_{\mu} + e^{+}$$

$$e^{+} + e^{-} \rightarrow \nu_{\mu} + \bar{\nu}_{\mu}$$

$$p + p \rightarrow \Sigma^{+} + \pi^{+}$$

$$\Sigma^{\circ} \rightarrow \Lambda^{\circ} + \gamma$$

$$D^{-} \rightarrow K^{+} + \pi^{-} + \pi^{-}$$

$$\Lambda^{\circ} + p \rightarrow K^{-} + p + p$$

- 2) Upper bound on neutrino mass from supernova explosion SN1987A. On 23.2.1987 a supernova explosion was observed in the Great Magellan Cloud that is 1.54×10^{18} km = 1.63×10^5 light years away. So the event happened 1.63×10^5 years ago. Two neutrino detectors independently observed a short burst of neutrinos with energies between 35 and 13 MeV within a 10 second time interval. Assume the neutrinos all started out at the same moment. Show that for two neutrinos with energies E_1 and E_2 the difference in arrival times is

$$\Delta t = t_2 - t_1 \approx \frac{Lm^2}{2c} \left(\frac{1}{E_2^2} - \frac{1}{E_1^2} \right)$$

Use this expression and the data to derive an upper bound for the neutrino mass. $L = 1.5 \times 10^{18}$ km.

- 3) Neutron-Antineutron oscillation. If the baryon number is conserved, the transition $n \leftrightarrow \bar{n}$, known as “neutron oscillation” is forbidden. The experimental limit on the time scale of such oscillations in free space is $\tau_{n \leftrightarrow \bar{n}} \geq 3 \times 10^6$ s. Let H_0 be the Hamiltonian in the absence of any interaction which mixes n and \bar{n} . Then:

$$H_0 |n\rangle = m_n c^2 |n\rangle \quad , \quad H_0 |\bar{n}\rangle = m_n c^2 |\bar{n}\rangle$$

for states at rest. Let H' be the interaction which turns n into \bar{n} and vice versa:

$$H'|n\rangle = \varepsilon|\bar{n}\rangle \quad , \quad H'|\bar{n}\rangle = \varepsilon|n\rangle$$

where ε is real and H' does not flip the spin.

Start with a neutron at $t = 0$ and calculate the probability that it will be observed to be an antineutron at time t . When the probability is first equal to 50%, call the time $\tau_{n \leftrightarrow \bar{n}}$. In this way convert the experimental lower limit on $\tau_{n \leftrightarrow \bar{n}}$ into an upper limit on ε .