Freie Universität Berlin Fachbereich Physik May 16th 2017 Prof. Dr. Roland Netz Douwe Bonthuis Julian Kappler Philip Loche

Statistical Physics and Thermodynamics (SS 2017)

Problem sheet 5

Hand in: Friday, May 26 during the lecture

http://www.physik.fu-berlin.de/en/einrichtungen/ag/ag-netz/lehre/

1 Total differential (6 points)

The differential $\alpha(x,y) = A(x,y)dx + B(x,y)dy$ is called *exact* if there is a function F(x,y), so that $dF = \frac{\partial F}{\partial x}dx + \frac{\partial F}{\partial y}dx$ is equal to α .

a) Show if α is exact, then A(x, y) and B(x, y) satisfy the condition

$$\frac{\partial A}{\partial y} = \frac{\partial B}{\partial x}.\tag{1}$$

(1 point)

Remark: The converse only holds for simply connected regions. In general just because equation (1) holds this does not implies that there is a globally defined function F such that α is an exact differential ¹.

Consider $\alpha(x, y) := (\cos x - x \sin x + y^2) dx + 2xy dy$ and $\beta(x, y) := (x^2 - y) dx + x dy$

b) Do $\alpha(x, y)$ and $\beta(x, y)$ fulfill equation (1)? (2 points)

c) Are α and β exact? If yes find a function F(x, y) such that $dF = \alpha, \beta$ and F(0, 0) = 0. (2 points)

d) Find a curve (x, y(x)) in space where $F(x, y) = F(\pi, 1)$ is constant. Plot y(x) in the range $-2\pi < x < 2\pi$. (1 point)

2 Entropy of two subsystems (7 points)

Consider two subsystems with the number of microstates $\Gamma_1(U_1)$ and $\Gamma_2(U_2)$ and their energies U_1 and U_2 respectively. The subsystems can exchange energy so that the total energy of the system is $U = U_1 + U_2$.

a) Write down the number of microstates of the total system $\Gamma(U)$. Use that $\ln \Gamma(U) = S(U)/k_B$ to express the total entropy S(U) in terms of the entropies of the two subsystems $S_1(U_1)$ and $S_2(U_2)$. (1 point)

b) Calculate S(U) for the explicitly given entropies $S_1(U_1) = -a_1(U_1 - U_1^0)^2$ and $S_2(U_2) = -a_2(U_2 - U_2^0)^2$ with $a_1 > 0, a_2 > 0$. (3 points)

c) The entropy and the energy are extensive quantities, so they scale with the number of particles N. What does this imply for the scaling of a_1 and a_2 ? (1 point)

- d) Using your results from c), how do the terms in S(U) scale? Is your result extensive for large N. (1 point)
- e) What is the temperature T of both systems and how does this scale with the number of particles N? (1 point)

¹https://en.wikipedia.org/wiki/Closed_and_exact_differential_forms

3 Scale heights (7 points)

In the following you will derive the "scale height", an estimate for the atmospheric height of a planet. Assume that all particles in the atmosphere got the mass m and are in thermal equilibrium at temperature T regardless of the their height h above the planet's surface. The particles are subjected to a gravitational potential with the energy U = -GmM/(R+h), with G the gravitation constant, M the planet's mass and R its radius.

a) Expand U for small h up to the linear order and write down the non normalized probability distribution $\rho(h)$ for the approximated potential. Use $g = GM/R^2$ as an abbreviation for the gravitational acceleration. Argue up to which height your approximation is valid. (2 points)

Hint: For the validity of your expansion consider higher order terms.

b) Normalize the distribution in terms of k_B , T, m and g for the approximated potential. Is the exact distribution normalizable? If not, what does this mean physically? (2 points)

c) Calculate the mean (scale height) and its standard deviation (2 points)

The following table lists some characteristics of Earth and Mars

	Earth	Mars
radius (R)	$6371 \ km$	$3390 \ km$
gravitational acceleration (g)	$9.82 \ m/s^2$	$3.74 \ m/s^2$
average temperature (T)	287 K	226 K
mass of atmospheres particles (m)	$28 \ u \ (N_2)$	$44,01 \ u \ (CO_2)$

d) Calculate the atmospheric scale heights and its standard deviation for Earth and Mars ($k_B = 1.38065 \times 10^{-23} J/K$). Is your approximation of a) still valid? (1 point)